

A HOT-ELECTRON FAR-INFRARED DIRECT DETECTOR

B.S. Karasik, W.R. McGrath, and H.G. LeDuc

*Center for Space Microelectronics Technology, Jet Propulsion Laboratory,
California Institute of Technology, Pasadena, CA 91109*

A new approach is proposed to improve the sensitivity of direct-detection bolometers at millimeter, submillimeter and far-infrared wavelengths. The idea is to adjust a speed of the thermal relaxation of hot-electrons in a nanometer size normal metal or superconductive transition edge bolometer by controlling the elastic electron mean free path. If the bolometer contacts are made of a superconductor with high critical temperature (Nb, Pb etc.) then the thermal diffusion into the contacts is absent because of the Andreev's reflection and the electron-phonon relaxation is the only mechanism for heat removal. The relaxation rate should behave as T^4l at subkelvin temperatures (l is the electron elastic mean free path) and can be reduced by factor of 10-100 by decreasing l . Then an antenna- or waveguide-coupled bolometer with a time constant $\sim 10^{-3}$ to 10^{-5} s at $T \approx 0.1$ -0.3 K will exhibit photon-noise limited performance in millimeter and submillimeter range. The choice of the bolometer material is a tradeoff between a low electron heat capacity and fabrication.

A state-of-the-art bolometer currently offers $NEP \approx 10^{-17}$ W $\sqrt{\text{Hz}}$ at 100 mK along with a ≈ 2 msec time constant. The bolometer we propose will have a figure-of-merit, $NEP\sqrt{\tau}$, which is 10^3 times smaller. This will allow for a tremendous increase in speed which will have a significant impact for observational mapping applications. Alternatively, the bolometer could operate at higher temperature with still superior sensitivity. This device can significantly increase a science return and reduce the cost for future observational missions.

This research was performed by the Center for Space Microelectronics Technology, Jet Propulsion Laboratory, California Institute of Technology, and was sponsored by NASA, Office of Space Science.